# Torsional oscillations of the large-scale circulation in turbulent Rayleigh-Bénard convection

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Measurements over the Rayleigh-number range  $10^8 \lesssim R \lesssim 10^{11}$  and Prandtl-number range  $4.4 \le \sigma \le 29$  that determine the torsional nature and amplitude of the oscillatory mode of the large-scale circulation (LSC) of turbulent Rayleigh-Bénard convection are presented. For cylindrical samples of aspect ratio  $\Gamma = 1$  the mode consists of an azimuthal twist of the near-vertical LSC circulation plane, with the top and bottom halves of the plane oscillating out of phase by half a cycle. The data for  $\Gamma = 1$  and  $\sigma = 4.4$  showed that the oscillation amplitude varied irregularly in time, yielding a Gaussian probability distribution centred at zero for the displacement angle. This result can be described well by the equation of motion of a stochastically driven damped harmonic oscillator. It suggests that the existence of the oscillations is a consequence of the stochastic driving by the small-scale turbulent background fluctuations of the system, rather than a consequence of a Hopf bifurcation of the deterministic system. The power spectrum of the LSC orientation had a peak at finite frequency with a quality factor  $Q \simeq 5$ , nearly independent of R. For samples with  $\Gamma \geqslant 2$  we did not find this mode, but there remained a characteristic periodic signal that was detectable in the area density  $\rho_p$  of the plumes above the bottom-plate centre. Measurements of  $\rho_p$  revealed a strong dependence on the Rayleigh number R, and on the aspect ratio  $\Gamma$  that could be represented by  $\rho_p \sim \Gamma^{2.7\pm0.3}$ . Movies are available with the online version of the paper.

#### 1. Introduction

Understanding turbulent Rayleigh–Bénard convection (RBC) in a fluid heated from below (see, for instance, Siggia 1994; Kadanoff 2001; Ahlers, Grossmann & Lohse 2002) remains one of the challenging problems in nonlinear physics. It is well established that a major component of the dynamics of this system is a large-scale circulation (LSC) (Sano, Wu & Libchaber 1989; Castaing et al. 1989; Ciliberto, Cioni & Laroche 1996; Qiu & Tong 2001a; Funfschilling & Ahlers 2004; Sun, Xia & Tong 2005b; Tsuji et al. 2005). The LSC plays an important role in many natural phenomena, including atmospheric and oceanic convection (see, for instance, van Doorn et al. 2000) where it has an impact on climate, and convection in the outer core of Earth (see, for instance, Glatzmaier, Coe & Roberts 1999) where it is responsible for the generation of the magnetic field.

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In this paper we consider cylindrical samples, primarily with aspect ratio  $\Gamma \equiv D/L \simeq 1$  (D is the diameter and L the height). For these the LSC consists of a single convection roll, with both downflow and upflow near the sidewall but at azimuthal locations  $\theta$  that differ by  $\pi$ . An interesting property of the LSC is a twisting azimuthal oscillation mode with frequency  $f_0 = \omega_0/2\pi$ , where  $\omega_0$  is the angular oscillation frequency. The spatial nature of this mode has become apparent only recently (Funfschilling & Ahlers 2004; Resagk *et al.* 2006; Xi, Zhou & Xia 2006). It is a time-periodic twist of the circulation plane of the LSC that consists, at a given moment, of a rotation in opposite azimuthal directions in the top and the bottom half of the sample (Funfschilling & Ahlers 2004). In the present paper the existence of this mode is confirmed by quite different measurements, its spatio-temporal structure is further illuminated, and measurements of its properties are extended to a larger Rayleigh- and Prandtl-number range.

Although the geometrical features of this mode were unclear until recently, a characteristic frequency was measured much earlier in a number of single-point determinations of the temperature or the velocity (Heslot, Castaing & Libchaber 1987; Castaing et al. 1989; Ciliberto et al. 1996; Takeshita et al. 1996; Cioni, Ciliberto & Sommeria 1997; Qiu, Yao & Tong 2000; Qiu & Tong 2001a, b; Niemela et al. 2001; Qiu & Tong 2002; Qiu et al. 2004). Depending on the nature of the probe, and the probe location within the sample, such measurements could determine either the inverse turnover time  $1/\mathcal{T}$  of the LSC or its azimuthal oscillation frequency  $f_0$ . In either case, it was found that, within experimental resolution,  $f_0$  is equal to  $1/\mathcal{T}$  over a wide parameter range (Qiu & Tong 2002; Funfschilling & Ahlers 2004; Brown, Funfschilling & Ahlers 2007). The reason for this equality is not known at this time.

Consistent with the twisting nature of the mode, we detected no oscillations in the orientation  $\theta_0$  of the circulation plane at the horizontal mid-plane of the sample. Instead,  $\theta_0$  revealed a random time dependence that could be described well by azimuthal diffusion (Brown, Nikolaenko & Ahlers 2005a; Sun, Xi & Xia 2005a; Brown & Ahlers 2006b; Xi et al. 2006), interrupted occasionally by larger re-orientations, as discussed before (Brown et al. 2005a; Brown & Ahlers 2006b; Xi et al. 2006). The out-of-phase oscillations were found in the angles  $\theta_t$  and  $\theta_b$ measured at vertical positions L/4 above and below the mid-plane respectively. Whereas a purely periodic oscillation with characteristic amplitude A should yield probability-distribution functions  $p_{\theta_b'}$  of  $\theta_b' = \theta_b - \theta_0$  and  $p_{\theta_t'}$  of  $\theta_t' = \theta_t - \theta_0$  with pronounced peaks at angles close to  $\pm A$ , we found that  $p_{\theta_h}$  and  $p_{\theta_l}$  were Gaussian distributed about  $\theta_b' = 0$  or  $\theta_t' = 0$ . Such a distribution can be reproduced by a damped harmonic oscillator driven by white noise (see, for instance, Gitterman 2005). Thus the measurements suggest that the existence of the oscillatory mode is a consequence of the stochastic driving by the small-scale turbulent background fluctuations, and that it is not a consequence of a Hopf bifurcation of the deterministic system.

Consistent with the conclusions of the previous paragraph, we found that the oscillatory mode has oscillating but gradually decaying time-correlation functions and a Lorentzian peak in the height-dependent power spectrum (or structure factor) S(z, f) of the LSC orientation  $\theta(z, t)$ . For  $\Gamma = 1$  and Prandtl number  $\sigma = 4.4$  we present results for the height  $S_0$ , centre frequency  $f_0$ , half-width  $\sigma_f$ , and total power  $P = \pi \sigma_f S_0$  of this peak as a function of the Rayleigh number

$$R = \frac{\alpha g \Delta T L^3}{\kappa \nu} \tag{1.1}$$

( $\alpha$  is the isobaric thermal expansion coefficient, g the acceleration due to gravity, and  $\Delta T$  the applied temperature difference). Although the half-width increases with R, the centre frequency increases also with the result that the quality factor  $Q = f_0/(2\sigma_f) \simeq 5$  remains nearly independent of R. This suggests that the oscillations will survive up to values of R well above those available in the laboratory, unless a dramatic event (such as the disappearance of the LSC itself) intervenes and invalidates a simple power-law extrapolation of the R-dependences of  $f_0$  and  $\sigma_f$ .

We find that the oscillatory mode exists only for aspect ratios near one. For larger  $\Gamma$  we do not find the azimuthal oscillation, but instead observe that the plume area density above the centre portion of the bottom plate oscillates periodically. Thus, as the aspect ratio changes, the nature of the time-periodic mode of the LSC changes its character.

#### 2. Experimental apparatus and methods

#### 2.1. The samples

Measurements using water at a mean temperature of 40°C (Prandtl number  $\sigma \equiv v/\kappa = 4.38$  and  $v = 6.7 \times 10^{-7}$  m² s<sup>-1</sup>) as the fluid were made for the two cylindrical samples with  $\Gamma \simeq 1$  described by Brown *et al.* (2005*b*) as the 'medium' and the 'large' sample. The top and bottom plates of both were made of copper. The Plexiglas sidewall had a thickness of 0.32 (0.63) cm for the medium (large) sample. The samples had L(D) = 24.76 (24.81) and 50.61 (49.69) cm respectively. Rayleigh numbers were determined from temperature measurements with five thermometers imbedded in each of the top and bottom plates.

Measurements using methanol with  $\sigma=6.0$  and  $\nu=5.79\times10^{-7}\,\mathrm{m^2\,s^{-1}}$  or 2-propanol with  $\sigma=28.9$  and  $\nu=1.77\times10^{-6}\,\mathrm{m^2\,s^{-1}}$  as the fluid were made in three smaller cylindrical samples. They each had a 6.35 mm thick aluminium bottom plate with a mirror top surface and a 3.15 mm thick sapphire top plate. A high-density polyethylene sidewall was sealed to the top and bottom plates by ethylenepropylene O-rings. One of the smaller samples had  $\Gamma\simeq1$  with L (D)=8.74 (8.66) and had been used before by Xu *et al.* (2000) and by Funfschilling & Ahlers (2004). We shall call it the 'small' sample. The other two had L (D)=4.48 (8.97) and 3.00 (8.96) cm respectively. They will be referred to as the  $\Gamma\simeq2$  and  $\Gamma\simeq3$  sample respectively. The bottom-plate temperature was determined with several thermistors imbedded in the plate. The top temperature was inferred from measurements of the temperature of the bath that circulated above the top sapphire plate.

#### 2.2. Sidewall-temperature measurements

For the medium and large sample, three sets of eight thermistors each, equally spaced around the circumference at the three vertical positions -L/4, 0, and L/4 (we take the origin of the vertical axis at the horizontal mid-plane of the sample) and labelled  $i=0,\ldots,7$ , etc. as shown in figure 1, were imbedded in small holes drilled horizontally into but not penetrating the sidewall. The thermistors were able to sense the adjacent fluid temperature without interfering with delicate fluid-flow structures. We measured the temperature of each thermistor with a sampling period of about 2.5 s. Since the LSC carried warm (cold) fluid from the bottom (top) plate up (down) the sidewall, these thermistors detected the location of the upflow (downflow) of the LSC by indicating a relatively high (low) temperature.

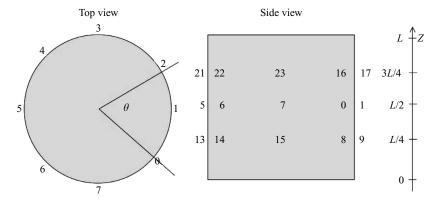


FIGURE 1. The location of the sidewall thermometers. For the top view only the thermometer locations for the set located a distance L/2 above the bottom plate (on the horizontal mid-plane, or z = 0) are shown. The side view indicates the location of all three sets.

To determine the orientation and strength of the LSC, we fitted the function

$$T_i = T_0 + \delta \cos\left(\frac{i\pi}{4} - \theta_0\right), \ i = 0, ..., 7,$$
 (2.1)

separately at each time step, to the eight temperature readings obtained from the thermistors at height 0. The fit parameter  $\delta$  is a measure of the temperature amplitude of the LSC and  $\theta_0$  is the azimuthal orientation of the plane of the LSC circulation. As defined here, the orientation  $\theta_0$  is on the side of the sample where the LSC is warm and upflowing and is measured relative to the location of thermometer zero. We calculated orientations  $\theta_t$  and  $\theta_b$  and amplitudes  $\delta_t$  and  $\delta_b$  for the top and bottom levels at L/4 and -L/4 separately by the same method as for the middle row. Results from this method of determining the orientation and strength of the LSC have been reported in several previous publications (Brown *et al.* 2005*a*; Ahlers, Brown & Nikolaenko 2005; Brown & Ahlers 2006*a*, *b*)

#### 2.3. Shadowgraph measurements

For the small,  $\Gamma \simeq 2$  and  $\Gamma \simeq 3$  samples the LSC movement and plume area density close to but above (below) the bottom (top) plate were examined by means of shadowgraph images (de Bruyn *et al.* 1996) as described in part already by Funfschilling & Ahlers (2004). Figure 2(a) shows a typical image obtained for the small sample. This image had a background image taken with  $\Delta T = 0$  subtracted, and the ratio was re-scaled to cover a suitable grey range. An elongated narrow dark stripe such as the one near the upper right of the image corresponds to a stripe of relatively warm fluid. For most of our work the optics was adjusted so as to emphasize the relatively warm plumes near the bottom plate. Cold plumes near the top plate are also discernible in the figure, but provide a relatively weaker bright signal. The two types of plumes are readily separated by image processing.

When the sample axis is parallel to the gravitational acceleration, the LSC orientation  $\theta_0$  meanders nearly randomly in the azimuthal direction (Sun *et al.* 2005a, Brown & Ahlers 2006a, b; Xi *et al.* 2006) on time scales much longer than the oscillation period. Some measurements were made under these conditions, but for others the sample was tilted by an angle  $\beta \simeq 1^{\circ}$  relative to gravity. This led to a strongly preferred azimuthal orientation of the LSC circulation-plane (Ahlers *et al.* (2005); Brown *et al.* (2005a); Brown & Ahlers 2006b). In figure 3 the coordinate

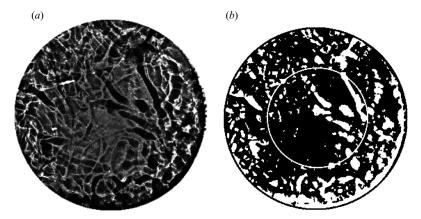


FIGURE 2. (a) Background-image-divided shadowgraph image of the small sample ( $\Gamma = 1$ ) with 2-propanol at a mean temperature of 40°C. The Rayleigh number was  $R = 7.8 \times 10^8$ . (b) A binary rendering of the image in (a) obtained by applying a threshold and inverting the grey scale. The white circle illustrates the area used for the plume area-density measurements.

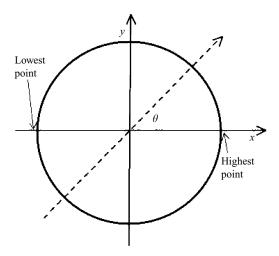


FIGURE 3. Schematic illustration of the coordinates used when the sample was tilted by an angle  $\beta \simeq 1^{\circ}$  relative to gravity.

system relative to the tilt direction is illustrated. For the small tilt angle used the shadowgraph measurements did not resolve any influence of the tilt on the properties of the oscillatory mode.

# 2.3.1. *Plume-velocity measurements*

For the small sample  $(\Gamma = 1)$  the speed  $s_p$  and direction  $\theta_p$  of the plume motion was determined from spatial cross-correlation functions

$$C(\delta \mathbf{r}, t) = \langle I(\mathbf{r}, t)I(\mathbf{r} + \delta \mathbf{r}, t + \delta t) \rangle$$

between two images separated in time by an interval  $\delta t$ . Here r is the two-dimensional position vector, t the time, and  $\langle \ldots \rangle$  indicates a spatial average. The method is illustrated in figure 4. Two images separated in time by  $\delta t = 0.9$  s (a and b) are shown. Each contains a dominant highly elongated plume. To a good approximation the plume moves in the direction of its long axis. This alignment of plumes by a shear

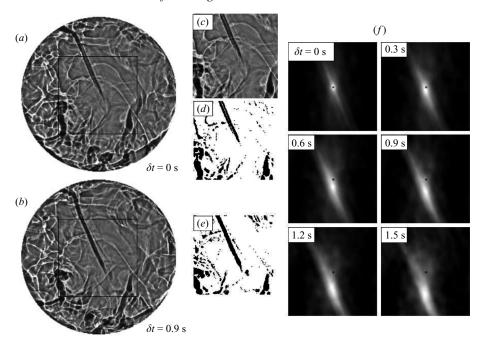


FIGURE 4. (a,b) Two background subtracted and re-scaled shadowgraph images separated in time by  $\delta t = 0.9$  s. (c) A central square section of (a). (d,e) One-bit rendering of the square sections from (a) and (b). (f) Spatial cross-correlation functions between images such as (d) and (e), with several different time delays  $\delta t$  ranging from zero to 1.5 s (the origin is indicated by a small black dot). The fluid was methanol at a mean temperature of  $40^{\circ}$ C and  $R = 4.6 \times 10^{8}$ .

flow is a well-known phenomenon and consistent with stability analysis for convection rolls in the presence of shear (Kelly 1994).

A central square (figure 4c) was extracted and converted to a one-bit image showing the plumes as black and most of the rest of the sample as white (figure 4d, e). The cross-correlation function between two such central squares, separated in time by various values indicated in the figure, are shown in figure 4(f). For  $\delta t = 0$  s there is the auto-correlation function which has a maximum at r = 0. Various time delays give various displacements of this peak, by an amount  $\delta r$  which, together with  $\delta t$ , gives the plume velocity  $v_p$ . The magnitude of  $v_p$  gives the plume speed  $s_p$ , and its direction is taken as the direction  $\theta_p$  of the plume motion. Indeed,  $\theta_p$  is seen to be close to the orientation of the long axis of the plume. This long-axis orientation tended to fluctuate about  $\theta_p$ . It is assumed that the LSC orientation is the direction of the plume motion. Time series typically contained from 8200 to 20200 images.

# 2.3.2. Plume area-density measurements

In order to determine the area density of plume  $\rho_p$ , a threshold equal to half the average intensity was applied to the background-subtracted images and all pixels with values above (below) the threshold value were set to zero (one). This isolated the warm plumes near the bottom plate, and the fraction of the pixels in the plume images (i.e. the fractional area occupied by their shadowgraph images) could be determined. This density was calculated using a central circular area of radius about half that of the cell as shown in figure 2(b), but the results were not very sensitive to the precise area chosen provided the annulus very near the sidewall (where there

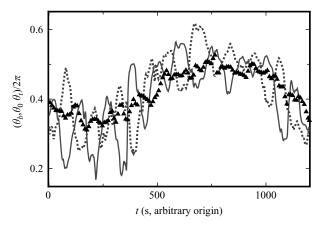


FIGURE 5. The large-scale-circulation orientations  $\theta_b$ ,  $\theta_0$ , and  $\theta_t$  calculated separately for the 3 sets of thermistors at  $R=3\times 10^{10}$  in the large sample. Dotted line:  $\theta_b$  from the bottom plane at a distance L/4 above the sample bottom (z=-L/4). Solid triangles:  $\theta_0$  from the mid-plane at a distance L/2 above the sample bottom (z=0). Solid line:  $\theta_t$  from the top plane at a distance 3L/4 above the sample bottom (z=L/4). One can see that the top- and bottom-plane orientations oscillate, out of phase with each other, around the mid-plane orientation.

is an accumulation of many rising plumes) was excluded. Although  $\rho_p$  depended somewhat on the intensity cutoff, its time and Rayleigh-number dependence and its probability-distribution function showed no significant change over a wide cutoff range. As defined here,  $\rho_p$  can change either because plumes move into or out of the measurement area, or because the rate of plume generation changes; the measurement does not distinguish between these two processes.

# 3. Experimental results for aspect ratio $\Gamma = 1$

# 3.1. Medium and large sample

In the small sample, the frequency  $f_0$  of temporal oscillations in the azimuthal direction of the motion of plumes across the bottom plate was previously observed using shadowgraphs (Funfschilling & Ahlers 2004). We can observe the same oscillations in the medium and large samples by looking at time series of the LSC orientations  $\theta_t$ ,  $\theta_0$ , and  $\theta_b$  at different heights. A short section of a much longer time series at the three heights is shown in figure 5 for  $R = 3 \times 10^{10}$  in the large sample. While the orientation given by each data set contains erratic motion, there is a tendency for the orientations at the top and bottom planes to oscillate out of phase with each other around the mid-plane orientation.

Power spectra of the three time series for  $\theta_i(t)$ , i=t,0, b are shown in figure 6. Peaks representing oscillations are clearly seen for the top  $(\theta_t)$  and bottom  $(\theta_b)$  orientations, but for the mid-plane  $(\theta_0)$  the spectrum gives no hint of oscillations. It is interesting to note that even a small misalignment of the sample relative to gravity, by  $10^{-2}$  rad or less, will lead to oscillations also at the mid-plane. In several experiments by others the sample was deliberately tilted slightly in order to limit the diffusion of the LSC orientation to a small angular range, and caution is in order when results for the torsional mode are compared with measurements under those conditions. We have carried out an extensive study of the influence of a tilt angle on the LSC dynamics, and will report on this in a separate publication.

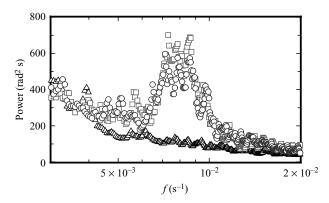


FIGURE 6. The smoothed power spectra of the three angles  $\theta_i(t)$  for  $R = 3 \times 10^{10}$  from the large sample. Triangles:  $\theta_0$ . Circles:  $\theta_t$ . Squares:  $\theta_b$ .

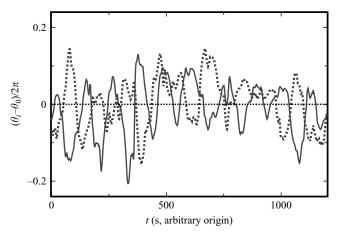


FIGURE 7. The large-scale-circulation orientations with the meandering of  $\theta_0$  subtracted out at  $R = 3 \times 10^{10}$  in the large sample. Dotted line:  $\theta_b - \theta_0$  from the bottom plane. Solid line:  $\theta_t - \theta_0$  from the top plane. One can see that the top- and bottom-plane orientations oscillate, out of phase with each other, around the mid-plane orientation.

The oscillations of  $\theta_b$  and  $\theta_t$  about  $\theta_0$  are shown more clearly by the deviations  $\theta_t - \theta_0$  and  $\theta_b - \theta_0$  of the top- and bottom-plane orientations from the mid-plane orientation; these are shown in figure 7 for the data of figure 5. To reveal the nature of the oscillations more quantitatively, we compute the cross-correlation functions

$$g_{i,j}(\tau) = \langle [\theta_i(t) - \theta_0(t)] \times [\theta_j(t+\tau) - \theta_0(t+\tau)] \rangle$$
(3.1)

where i and j can each refer to either the top- or bottom-plane orientations  $\theta_t$  or  $\theta_b$ , and where  $\langle ... \rangle$  without a subscript represents a time average. Normalized by  $\sqrt{g_{i,i}(0)g_{j,j}(0)}$ , these are shown in figure 8 for  $R=3\times 10^{10}$  and the large sample. The oscillations in the auto-correlation functions confirm that the top- and bottom-plane orientations oscillate, and the lack of a large background in the auto-correlations confirms that the oscillation is really around the mid-plane orientation  $\theta_0$  and not around some fixed orientation. The cross-correlation similarly shows oscillations, and the minimum of  $g_{t,b}(\tau)$  at  $\tau=0$  indicates that the top and bottom row oscillations are out of phase with each other by half a period. The oscillations shown here lead

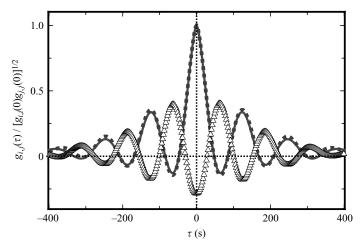


FIGURE 8. Normalized correlation functions  $g_{i,j}(\tau)/\sqrt{g_{i,i}(0)g_{j,j}(0)}$  for the top-, mid-, and bottom-row LSC orientations for  $R=3\times 10^{10}$ . Open triangles: cross-correlation between top and bottom rows (i=t,j=b). Solid dots: auto-correlation of bottom row (i=j=b). Solid line: auto-correlation of top row (i=j=t). The oscillations are very clear, with the minimum of  $g_{t,b}(\tau)$  at  $\tau=0$  indicating that the top- and bottom-plane orientations oscillate out of phase with each other.

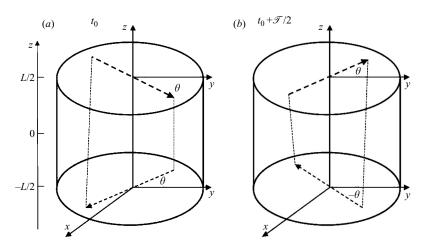


FIGURE 9. Schematic diagram of the LSC oscillation at time  $t_0$  when the oscillation is at an extremum (a), and at time  $t_0 + \mathcal{F}/2$  (i.e. half an oscillation period later) (b).

to the same picture of an azimuthal twisting, or torsional, oscillation of the LSC as found by Funfschilling & Ahlers (2004).

In figure 9 we show a schematic diagram that illustrates the nature of the oscillating mode, with the top and bottom out of phase. This mode can be described by

$$\theta(z,t) = \theta_0(t) + f(z)\theta'(t) \tag{3.2}$$

where  $\theta(z,t)$  is the height-dependent orientation of the circulation plane of the LSC. If we choose the origin of the vertical axis at the horizontal mid-plane as shown in the figure, then the oscillation amplitude becomes an odd function of z, i.e. f(z) = -f(-z). We set f(L/4) = 1 so that  $\theta'(t) = \theta_i(t) - \theta_0(t)$  as shown in figure 7.

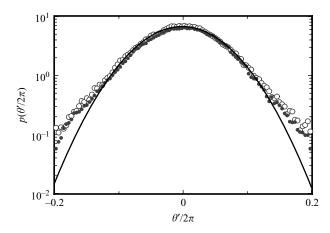


FIGURE 10. The probability distribution  $p_{\theta'}$  for  $\theta' = \theta_b - \theta_0$  (solid circles) and for  $\theta' = \theta_t - \theta_0$  (open circles) for  $R = 3 \times 10^{10}$  in the large sample. The solid line is a fit of a Gaussian function to the data near the peak. The fit gave a square root of the variance equal to 0.057, corresponding to about  $20^{\circ}$ .

Elsewhere we have shown that the background dynamics given by  $\theta_0(t)$  is primarily diffusive, albeit interrupted by various relatively sudden re-orientation events (Brown et al. 2005a; Brown & Ahlers 2006b). The twist angle  $\theta'(t)$  also has an irregular time dependence. In figure 10 we show the probability distribution  $p_{\theta'}$  as a function of  $\theta' = \theta_i - \theta_0$  with i = t, b. As shown by the solid line in the figure, the peak of  $p_{\theta'}$  can be fitted well by a Gaussian function. Further away from the peak there are deviations from the Gaussian form, with the experimental data yielding a somewhat larger  $p_{\theta'}$ . If the twist could be represented by a simple oscillation with  $\theta'(t) = A\cos(\omega t)$  with a constant (time-independent) A, then one would expect  $p_{\theta'}$  to have maxima at  $\theta' = \pm A$ . In contrast to this, the measurements indicate a peak at  $\theta' = 0$ . This is consistent with a time-dependent A(t) with a broad probability distribution  $p_A$  of A.

Recently Xi et al. (2006) studied the oscillations of the LSC circulation plane by particle-image velocimetry measurements in a horizontal plane located very near the top plate of their cylindrical sample with  $\Gamma = 1$ . They measured the peak-to-peak amplitude 2A(t) and found its average to be in the range from 35° to 55° with no clear dependence on R. Although they were not able to see the torsional nature of this mode by measurements at a single vertical position, they found a Gaussian-distributed A(t). Their results for the time average  $\langle A \rangle \simeq 22^{\circ} \pm 5^{\circ}$  of A are at least roughly consistent with our result for the square root of the variance of the azimuthal displacement  $\theta'$ , which we typically found to be near  $20^{\circ}$ .

A simple model for  $\theta'(t)$  that reproduces a Gaussian  $p_{\theta'}$  is given by the stochastic differential equation describing a damped driven harmonic oscillator

$$\ddot{\theta}' = -b\dot{\theta}' - \omega_0^2 \theta' + g(t), \tag{3.3}$$

where b is a damping coefficient,  $\omega_0$  is the known angular oscillation frequency, and g(t) is a Gaussian-distributed white-noise term. This equation produces oscillatory motion in appropriate parameter ranges, and it is known that diffusion in a harmonic well produces a Gaussian position distribution (Gitterman 2005). While we have no physical reason for using the model (3.3), it is similar in mathematical form to others used to describe the LSC dynamics (Brown & Ahlers 2006a, 2007, 2008). The orientation of the LSC has been found to have characteristics of diffusive meandering

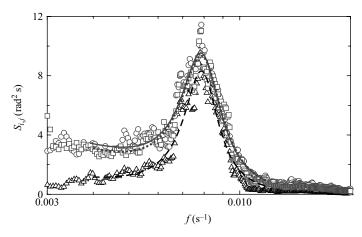


FIGURE 11. The smoothed power spectra  $S_{i,j}(f)$  of  $\theta'$  obtained from the Fourier transforms of each of the three correlation functions  $g_{i,j}(\tau)$  for  $R=3\times 10^{10}$  from the large sample. Triangles:  $S_{t,b}$ . Circles:  $S_{t,t}$ . Squares:  $S_{b,b}$ . Lines: a Lorentzian function with a exponential background is fit to the peak for  $S_{t,b}$  (dashed line),  $S_{t,t}$  (solid line), and  $S_{b,b}$  (dotted line).

(consistent with Gaussian-distributed white noise for g(t)) both at the mid-plane (Brown & Ahlers 2006a; Sun et al. 2005a) and near the top plate (Xi et al. 2006). The good agreement between the predictions of the model (3.3) and the data for  $p_{\theta'}$  indicates that the existence of the oscillatory mode is a result of the stochastic driving which we attribute to the small-scale turbulent fluctuations, and that this mode would decay without the driving.

The random time dependence of  $\theta'$  also leads to the decay of the correlation functions shown in figure 8. In order to provide a quantitative characterization of  $\theta'(t)$ , we show a smoothed power spectrum (the square of the modulus of the Fourier transform, or the 'structure factor')  $S_{i,j}(f)$  of  $\theta_t - \theta_0$  and  $\theta_b - \theta_0$  in figure 11. In addition, the Fourier transform of  $g_{i,b}(t)$  is shown. Particularly for  $S_{i,j}(f)$  one sees that there remain small backgrounds that are monotonically decreasing with f, although these backgrounds are much smaller than those seen in figure 6 for the spectra of  $\theta_i$ . In addition to the background there is a strong peak at  $f = f_0$  for all three spectra. Its finite width is attributable to the distribution of A(t). We fitted an exponential background and a Lorentzian peak

$$S_{i,j}(f) = \frac{S_0}{1 + [(f - f_0)/\sigma_f]^2} + a \exp \frac{-f}{b}$$
(3.4)

to the data for  $S_{i,j}(f)$  over the range  $0.5f_0 < f < 1.5f_0$  with fitting parameters  $S_0$ ,  $f_0$ ,  $\sigma_f$ , a, and b. In figure 11 one sees that the Lorentzian shape gives a good representation of the data.

In figure 12 we show the Rayleigh-number dependence of the scaled (i.e. dimensionless) fit parameters  $S_0\nu/L^2$  (the peak height),  $\sigma_f L^2/\nu$  (the peak halfwidth), and  $P_L = \pi S_0 \sigma_f$  (the total power under the Lorentzian peak excluding the background). The fit parameters for the bottom-plane, the top-plane and the cross-correlation data were quite similar, and thus in the figure their averages are displayed. One sees that the width of the peak increases and the height decreases with increasing R. For the dimensionless height the data from the two samples do not collapse onto a single curve, as would have been expected because the two samples have the same

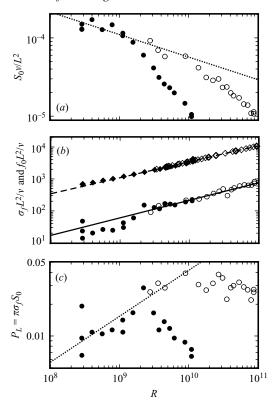


FIGURE 12. Fit parameters from (3.4) for the power spectra. These results are the averages obtained from the three spectra shown in figure 11. Solid symbols: medium sample. Open symbols: large sample. (a) Dimensionless peak height  $S_0\nu/L^2$ . The dotted line corresponds to a power law with an exponent of -0.45. (b) Dimensionless peak half-width  $\sigma_f L^2/\nu$  (circles) and centre frequency  $f_0$  (diamonds). The solid (dashed) line corresponds to a power law with an exponent of 0.55 (0.50). (c) The total power given by the area under the Lorentzian peak.

aspect ratio. It is not clear to us what physical phenomenon would produce this effect. Deviations from the Boussinesq approximation (Ahlers et al. 2006, 2007) which could influence the results at large R come to mind; however, for other properties such as the centre temperatures of the samples these effects come into play only at much larger values of the applied temperature differences. For the total power the R-dependences of  $S_0$  and of  $\sigma_f$  partly cancel, yielding a power with only a weak dependence on R. As was seen for  $S_0$ , the data for the two samples do not collapse. The dimensionless half-width and centre frequency each fall on the same curve for the two samples, as one would expect for samples of the same  $\Gamma$  and  $\sigma$ . The centre frequency is closely related to the inverse turnover time, and thus the Reynolds number, of the LSC. This is discussed in detail elsewhere (Brown et al. 2007).

A question of interest is whether the LSC and/or its oscillations disappear at large R. As well as looking at the total power of the oscillations, one might consider the mode to be unresolvable when the width  $2\sigma_f$  of the peak becomes significantly greater than the frequency  $f_0$  of the mode, i.e. when the quality factor  $Q = f_0/(2\sigma_f)$  becomes significantly less than one. This comparison can be made using the data for  $f_0$  and  $\sigma_f$  in figure 12(b). A power-law fit to the half-width for  $R > 3 \times 10^9$  yields  $\sigma_f \propto R^{0.55}$ ,

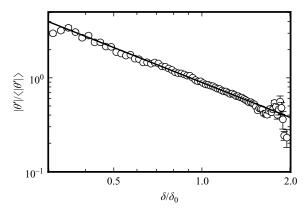


FIGURE 13. The time-average twist displacement  $\theta'$  as a function of  $\delta$  for  $R = 3 \times 10^{10}$  in the large sample. Solid line: power-law fit to the data.

whereas such a fit to  $f_0$  yields  $f_0 \propto R^{0.50}$ . In our range of R the half-width is about 1/10 of the frequency as seen in figure 11, i.e.  $Q \simeq 5$ . The closeness of the effective exponents for the two parameters implies that an extrapolation of the width to larger R remains small compared to the extrapolated frequency over a very large range of R, and thus that Q is nearly independent of R. This suggests that the torsional oscillation will not die out at any R obtainable in the laboratory owing to the spectral width becoming larger than the peak frequency unless the observed R-dependences change dramatically at higher R.

To gain some more insight into this mode, we bin the data based on the value of  $\delta$  and plot the average magnitude of the twist displacement  $|\theta'|$  as a function of  $\delta$  at the mid-plane in figure 13. A power-law fit over the range  $0.35 \le \delta/\langle \delta \rangle \le 1.9$  gives  $|\theta'| \propto \delta^{-1.24}$ . Since we do not find any corresponding oscillations in  $\delta$ , this relationship indicates that the amplitude of oscillations in  $\theta'$  decreases strongly with increasing  $\delta$ . It suggests that this scaling law is similar to one found for the rotation rate  $|\dot{\theta}_0|$  (Brown & Ahlers 2006b).

#### 3.2. Small sample

# 3.2.1. On the nature of plumes

The nature of the relatively warm (cold) localized volumes of fluid, also known as 'plumes', that are emitted by the bottom (top) boundary layer has been investigated numerous times, in most cases by using the shadowgraph method. This method detects the integral along the path of observation of the refractive-index variations orthogonal to that path. When viewed from above (Tanaka & Miyata 1980; Zocchi, Moses & Libchaber 1990; Gluckman, Willaime & Gollub 1993; Verzicco 2002; Funfschilling & Ahlers 2004; Haramina & Tilgner 2004; Puthenveettil & Arakeri 2005; Zhou, Sun & Xia 2007) plumes have generally given the appearance of structures that are localized in only one horizontal direction and extended in an orthogonal horizontal direction. These structures often are referred to as 'sheet' plumes, although this term may be misleading because their two-dimensional extent does not seem to be established. We prefer to call them 'line' plumes because it appears more likely to us that they are one-dimensional excitations in the marginally stable boundary layers.

Examples of plume visualizations taken from above have been shown in figures 2 and 4. Four additional examples are shown in figure 14 for several values of R, and movies showing the dynamics for two of these examples are available with the online

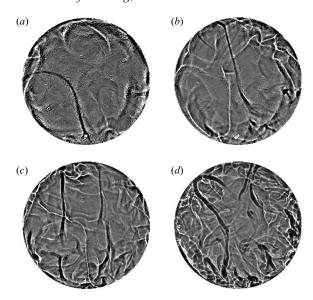


FIGURE 14. Shadowgraph images for the small  $\Gamma=1$  sample with methanol at a mean temperature of 40°C. (a)  $R=5.7\times10^7$ . (b)  $R=1.2\times10^8$ . (c)  $R=2.3\times10^8$ . (d)  $R=4.6\times10^8$ . Movies corresponding to (b) and (d) are available with the online version of the paper.

version of this paper. A plume near the bottom plate appears as a narrow elongated relatively dark stripe. Since, to lowest order, the shadowgraph method under certain conditions yields the vertical average of the horizontal Laplacian of the refractive-index (and thus of the temperature) field (Rasenat *et al.* 1989; de Bruyn *et al.* 1996; Trainoff & Cannell 2002), such a dark stripe is generally framed on each side by an even more narrow light stripe. This entire structure then corresponds to a line plume. Analogous bright lines flanked by dark ones correspond to cold line plumes near the top, but in our work these were made less prominent by appropriate adjustment of the optics. Since the structure is viewed from above, we can have no direct information about its vertical extent. Thus, it could indeed be a sheet-like structure with significant vertical extent, or it could be a line structure with a predominantly horizontal orientation.

However, these structures are generally not observed in shadowgraphs taken in the horizontal direction (Zocchi et al. 1990; Moses et al. 1991; Moses, Zocchi & Libchaber 1993; Ciliberto et al. 1996; Zhang, Childress & Libchaber 1997; Du & Tong 1998; Qiu & Tong 2001 b; Shang et al. 2003; Xi, Lam & Xia 2004), where plumes usually have the appearance of lines, often emanating from and extending away from the boundary layer and terminating in a mushroom-shaped cap. We shall refer to these structures as 'mushroom plumes'. Since line plumes do not appear in lateral observations, it appears that they are confined to the immediate vicinity of the bottom or top boundary layer where visualization from the side becomes difficult. The mushroom plumes are found in transverse observations primarily in the lower or upper portion of the sample, but often well separated from the bottom or top boundary layer. They are swept laterally by the LSC toward the sidewall. Near the sidewall they rise or fall owing to their buoyancy. From all of these observations the picture emerges that plumes are born as line excitations of the boundary layers, extended in one horizontal direction only and having initially very little if any vertical extent. These excitations are also swept laterally and in very close proximity to the

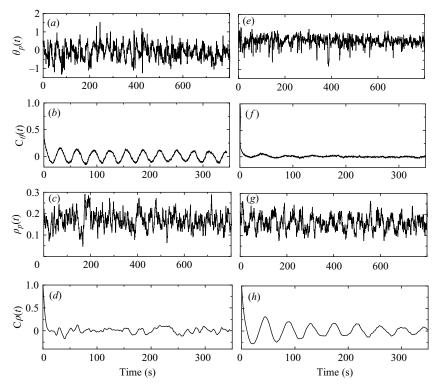


FIGURE 15. (a) The plume direction  $\theta_p$  and (c) density  $\rho_p$ , and their correlation functions (b)  $C_{\theta}(\tau)$  and (d)  $C_{\rho}(\tau)$ , as a function of time t or  $\tau$ :  $\Gamma = 1$  with 2-propanol and  $R = 7.8 \times 10^8$ . (e-h) corresponding plots for  $\Gamma = 2$  with 2-propanol and  $R = 1.5 \times 10^8$ .

boundary layer by the LSC. In the presence of the LSC they orient themselves with their long axis in the direction of the LSC. This alignment phenomenon is well known in other contexts, where convection rolls align their axes with a prevailing wind or in the direction of a shear (for a review, see for instance Kelly 1994). At some point along their path, and especially when impinging upon the sidewall, line plumes separate from the boundary layer and evolve into mushroom plumes with their characteristic mushroom-shaped heads. This process leaves the sample centre relatively free of plumes. Consequently observations from the top and not too close to the sidewall reveal primarily the line plumes very near the top and bottom boundaries.

# 3.2.2. Plume-direction oscillation and plume area density

As reported before from experiments using methanol (Funfschilling & Ahlers 2004), the azimuthal direction  $\theta_p(t)$  of the lateral motion of plumes near the top or bottom plate oscillates in time. This is illustrated in figures 15(a) and 15(b) for a Rayleigh number of  $7.8 \times 10^8$  using 2-propanol as the fluid. Figure 15(a) gives a short segment of a much longer time series of  $\theta_p(t)$ . The signal is quite noisy because occasionally there are too few or no plumes present and the calculated cross-correlation function between successive images (see figure 4) does not reflect plume motion. Although an algorithm could have been developed to eliminate these spurious points, it turned out that the time correlation-function

$$C_{\theta}(\tau) = \langle \theta_{n}(t)\theta_{n}(t+\tau) \rangle_{t} \tag{3.5}$$

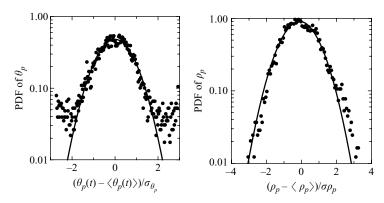


FIGURE 16. The probability-density functions (PDF) of  $\theta_p$  and  $\rho_p$  for  $\Gamma=1$  and 2-propanol. The horizontal axes were scaled by the square roots of the variances  $\sigma_{\theta_p}$  and  $\sigma_{\rho_p}$  of the data. The Rayleigh number was  $R=7.8\times10^8$ . The solid lines are fits of a Gaussian function to the data.

shown in figure 15(b) clearly reveals the oscillations and shows that they remain coherent over a long time period. We assume that the plume motion is slaved to the motion of the LSC and that  $\theta_p$  closely reflects the LSC flow direction  $\theta(\pm L/2, t)$  (see (3.2)) near the top and bottom plates. The same oscillations have been found also in velocity measurements using air (Resagk et al. 2006) as the fluid. Simultaneous measurements for methanol using the top (bright) and bottom (dark) plumes revealed that the oscillations near the top and bottom plate are out of phase with each other by  $\pi$  (Funfschilling & Ahlers 2004), consistent with the sidewall-temperature measurements reported above in § 3.1 and shown in figure 8.

The results for  $\theta_p$  lead to a power spectrum (not shown here) similar to those of  $\theta_t$  and  $\theta_b$  in figure 6, with a pronounced peak centred at a characteristic frequency  $f_0$  which corresponds to the inverse turnover time of the LSC. The Reynolds numbers corresponding to  $f_0$  already have been discussed by Brown *et al.* (2007).

The area density of plumes  $\rho_p(t)$  is illustrated in figure 15(c). One sees that  $\rho_p$  varied irregularly with time, with no obvious oscillatory component. This is borne out by the time auto-correlation function of  $\rho_p$ , which is shown in figure 15(d) and which meanders non-periodically about zero.

In figure 16 we show the probability-density functions (PDF) of  $\theta_p(t)$  and of the density  $\rho_p(t)$  for the example of figure 15. These functions show an unremarkable Gaussian distribution. Such a distribution has been noted (Zhou *et al.* 2007) with regard to other features of plumes in turbulent RBC.

The oscillation amplitudes of  $\theta_p$  can in principle be estimated from the power spectrum of  $\theta_p$  in a manner similar to that employed to obtain the data in figure 12 for  $\theta_t$  and  $\theta_b$ . However, the signal-to-noise ratio was considerably smaller for this method than it was for the sidewall-temperature measurements, and no definitive statement about the Rayleigh-number dependence of the amplitude could be made. Thus we only mention that typical values were in the range from 0.3 to 0.7 rad for  $5 \times 10^8 < R < 3 \times 10^9$ , both for 2-propanol and for methanol.

# 4. Results for aspect ratios $\Gamma > 1$

For the aspect ratios  $\Gamma = 2$  and  $\Gamma = 3$  the oscillations of  $\theta_b$  were very weak or absent entirely. This is illustrated in figure 15(e-h) for  $\Gamma = 2$  and 2-propanol. Figure 15(e)

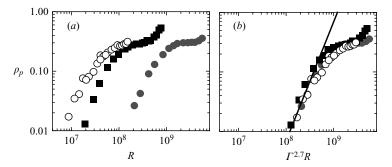


FIGURE 17. (a) The plume area-density above the bottom plate as a function of the Rayleigh number R. Solid circles:  $\Gamma = 1$ . Solid squares:  $\Gamma = 2$ . Open circles:  $\Gamma = 3$ . (b) The same data plotted as a function of  $\Gamma^{2.7}R$ . Solid line: a power law with an exponent of 2.0.

shows part of a time series of  $\theta_p$ . No oscillations are apparent to the eye. Figure 15(f) gives the corresponding correlation function. Whereas it shows a hint of oscillations at early times, these decay within a few cycles. This stands in stark contrast to the  $\Gamma=1$  case where  $C_\theta$  shows little or no decay over the time interval shown. On the other hand, for  $\Gamma=2$  the density of plumes (figure 15g) clearly oscillates. These ocillations are more clearly revealed by the correlation function in figure 15(h). This behaviour again stands in contrast to the  $\Gamma=1$  case where  $C_\rho$  showed no periodic component at all. The oscillations of  $\rho_p$  were found for all Rayleigh numbers above  $R\simeq 2\times 10^7$  for aspect ratio 2 and  $R\simeq 1.5\times 10^7$  for aspect ratio 3. The oscillation amplitudes of the densities were quite large (typically 50% of the mean value  $\langle \rho_p \rangle$ ). In spite of the difference in the temporal structure between  $\Gamma=1$  on the one hand and  $\Gamma=2$  and 3 on the other, the PDFs of both  $\theta_p$  and  $\rho_p$  were very similar for the two categories. We do not have any information about the geometrical features of the oscillating mode for the larger  $\Gamma$  cases. However, the difference presumably is associated with a major change in the structure of the flow.

Finally, in figure 17(a) we show the area density of the plumes above the bottom plate as a function of R for all three samples with 2-propanol. The data for the three aspect ratios fall on well separated curves. Interestingly, they essentially collapse onto a single curve when plotted against  $\Gamma^{2.7}R$ , revealing a strong  $\Gamma$ -dependence. A power-law scaling of  $\rho_p$  with R can be expected only when  $\rho_p \ll 1$  because the density has an upper limit of unity by virtue of its definition. The R-range of the data over which  $\rho_p \ll 1$  is rather small, and thus no effective exponent can be established with confidence. In addition the apparent plume width revealed by the shadowgraph method may depend on R. This effect would introduce an additional distortion of the measured R-dependence. As a guide to the R-dependence, the solid line in figure 17 is drawn with a slope of 2.

# 5. Summary and conclusions

In this paper we presented a detailed account of shadowgraph measurements, in part reported briefly before (Funfschilling & Ahlers 2004), of the azimuthal orientations  $\theta_p$  of plume motions across the top and bottom plates of turbulent samples of fluid in cylindrical containers and heated from below. These data are for aspect ratios  $\Gamma$  from 1 to 3. It is argued that  $\theta_p$  reflects the orientation of the LSC that is sweeping the plumes across the plates. For  $\Gamma=1$  the data revealed an azimuthal

LSC oscillation that had been shown before (Funfschilling & Ahlers 2004) to be a twisting mode of the LSC circulation plane, with the top and bottom halves oscillating out of phase by  $\pi$ . For  $\Gamma=2$  and 3 this mode was not present or extremely weak.

In addition to the results for  $\theta_p$ , shadowgraph measurements for the plume area density  $\rho_p$  near the centre of the bottom plate are reported. For  $\Gamma=1$ ,  $\rho_p$  varied irregularly in time but did not reveal the oscillations found in  $\theta_p$ . Interestingly, for the larger  $\Gamma=2$  and 3 (where  $\theta_p$  did not oscillate)  $\rho_p$  had a strong oscillating component. The geometrical nature of the mode leading to an oscillation of  $\rho_p$  and only irregular meandering of  $\theta_p$  is not yet revealed by the data.

For  $\Gamma=1$  the typical oscillation amplitude of  $\theta_p$  was in the range from 0.2 to 0.7 rad. For all three aspect ratios, the shadowgraph measurements revealed that both  $\rho_p$  and  $\theta_p$  had Gaussian probability distributions regardless of the absence or presence of oscillations. For small  $\rho_p$  (say  $\rho_p \lesssim 0.2$ ) the plume area density for all three  $\Gamma$  had a strong Rayleigh-number dependence, but at large R  $\rho_p$  saturated near  $\rho_p \simeq 0.4$  (we note that by virtue of its definition  $\rho_p < 1$ ). There also was a strong  $\Gamma$ -dependence of  $\rho_p(R)$ . The data for all three  $\Gamma$  could be collapsed onto a unique curve when plotted against  $\Gamma^{2.7\pm0.3}R$ .

We also presented new measurements of the LSC orientation  $\theta(z,t)$  by a very different method. These measurements are for two cylindrical samples with  $\Gamma=1$  over the Rayleigh-number range  $10^8 \lesssim R \lesssim 10^{11}$  and for  $\sigma=4.4$ . In this work we have determined  $\theta_b=\theta(-L/4,t),\ \theta_0=\theta(0,t),$  and  $\theta_t=\theta(+L/4,t)$  (we take the origin of the z-axis at the horizontal mid-plane) from measurements of the azimuthal dependence of the sidewall temperatures. The results confirm the shadowgraph observations in that they also show a torsional, i.e. twisting, mode of the LSC circulation plane, with  $\theta_b(t)$  and  $\theta_t(t)$  oscillating out of phase by  $\pi$  and  $\theta_0$  revealing no oscillatory contribution at all. Although each angle by itself shows a diffusive time evolution, with an oscillatory component added for  $\theta_t$  and  $\theta_b$ , the differences  $\theta_t'=\theta_t-\theta_0,\ t=b,t$ , reveal almost no meandering background at all, indicating that the oscillations are about  $\theta_0$  rather than about some other fixed angle.

For a purely sinusoidal oscillation of  $\theta'$  with characteristic amplitude A one expects a probability distribution  $p_{\theta'}$  of  $\theta'$  that is strongly peaked near  $\theta' = \pm A$ . Instead, we found that  $p_{\theta'}$  has a maximum at  $\theta' = 0$  and that it has a Gaussian shape when  $|\theta'|$ is not too large. This result is reproduced by the stochastic second-order differential equation that describes a damped harmonic oscillator driven by random noise. This equation for  $\theta'(t)$  is reminiscent of the equations that we discussed elsewhere (Brown & Ahlers 2006a, 2007, 2008) to explain the rapid re-orientations of the LSC orientation and the influence of Earth's Coriolis force on the LSC. In those equations the physical origins of all terms were clear, and the small-scale turbulent background was represented by a phenomenological additive noise term corresponding to experimentally determined diffusivities. In the present case we assume that the required stochastic term likewise represents the background fluctuations, but we do not have a physical explanation for the origin of the inertial term that is needed in order to reproduce the oscillations. An important conclusion to be drawn from  $p_{\theta'}$  is that the oscillations would not exist in the absence of the stochastic driving, and thus that they are not the result of a Hopf bifurcation of the deterministic system.

The oscillations of  $\theta_b$  and  $\theta_t$  lead to a peak at finite frequency of the power spectrum of  $\theta_i - \theta_0$  (i = t, b), with only a small broad background contribution. The peak could be described by a Lorentzian function. We presented data for the height  $S_0$ , half-width  $\sigma_f$ , centre frequency  $f_0$ , and total power  $P_L = \pi S_0 \sigma_f$  of this peak. We found that the

height  $S_0$  of the peak decreased and that the width  $\sigma_f$  increased with increasing R. Although both  $f_0$  and  $\sigma_f$  increase with R, the quality factor  $f_0/(2\sigma_f)$  is nearly independent of R, suggesting that the oscillatory mode will survive up to R-values well beyond those of the present data unless a dramatic change in the behaviour of the system intervenes.

An interesting question is how the oscillatory mode comes into existence as R increases from small values. We do not have data at sufficiently small R to provide an answer, but this issue was addressed by Qiu & Tong (2001 b) using measurements at  $\sigma = 5.4$ . From velocity and temperature data at single points in the sample, and correlation functions between them, they found a sharp transition at  $R_c \simeq 5 \times 10^7$  from a small-R state without coherent oscillations to another at larger R that exhibited the oscillations. At  $R_c$  the frequency observed by these authors had a finite value. Since we argued above that the oscillatory mode is not created via a Hopf bifurcation but instead is a consequence of the stochastic driving by the small-scale turbulent fluctuations, we interpret the observations of Qiu & Tong (2001 b) of a suden onset to imply that the driving begins abruptly as the Rayleigh number increases through  $R_c$ . Obviously, additional measurements will be required in order to clarify this important aspect of the LSC.

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